

# Japanese Balloon Telescope Program and Experiences

Yukihiro Takahashi  
*Department of Cosmo sciences,  
Hokkaido University*



# Situation of Venus exploration in Japan

Akatsuki (orbiter)

- waiting for 5-6 years

Akastuki-2 (orbiter)

- discussion started
- competitive to MELOS (Mars) and others

Balloon telescope

- (2009. Jun), 2012. Aug, 2013 summer, ...

Ground-based telescope

- 1.6 reflector (Pirka telescope in Hokkaido Univ.)

Funding situation in JAXA seems very serious...

# **Balloon Project Members**

## **Rikkyo Univ.**

Prof. Makoto Taguchi, PI

## **Tohoku Univ. (Engineering)**

Prof. Kazuya Yoshida

Asst. Prof. Yuji Sakamoto

## **ISAS/JAXA (Engineering)**

Dr. Yasuhiro Shoji

## **Hokkaido Univ. (Science and payload)**

Prof. Yukihiro Takahashi


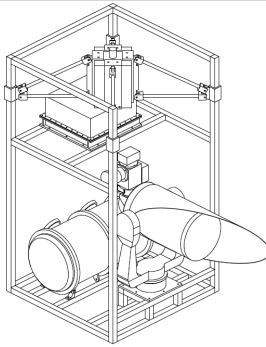
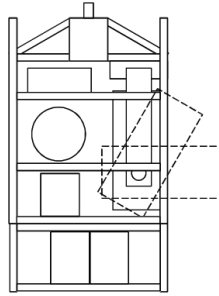
Asst. Prof. Makoto Watanabe

# Project Overview

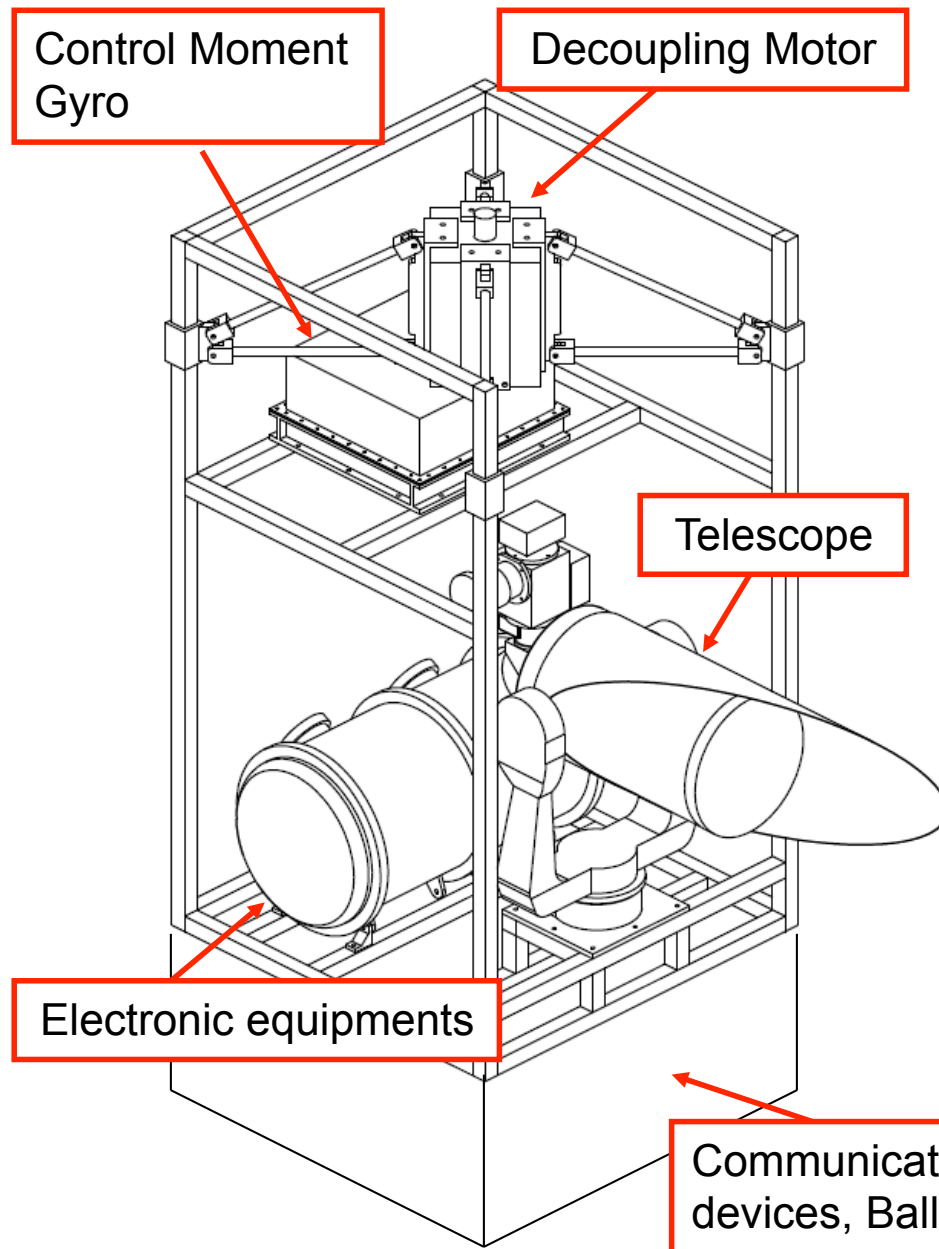
## Project aim :

To develop the technology of BBT gondola for the high resolution imaging of planets( the observation system, the pointing control system, and the others), and to conduct technology demonstrations in Japan.

- To conduct the flight operation for continuous observations in the polar stratosphere.

	2009, Hokkaido, Japan	2012, Hokkaido, Japan	2013~2014, polar region (TBD)
Gondola Overview			
Telescope	Schmidt Cassegrain, 300mmφ		Cassegrain-Nasmyth, 400mmφ
Purpose	- Demonstration and verification of the optical system and the pointing control system.	-Demonstration and verification of the optical system and the pointing control system. -Basic data collection for “multi-target observation”.	-Continuous observation for about 24~48 hours. -Multi-target observation -Basic data collection for longer duration observation.

# Gondola System (2012 Flight model)



**Size** : 1.0(W)×1.0(D)×2.5(H) m

**Weight** : Approx. 600 kg

**Power System** :

Solar Cell Panel

Li-ion Battery (25.9V, 50Ah)

**Data Communication** :

Downlink : 57600bps, telemetry data

Uplink : 1200bps, serial command

**Data Storage** :

On board memories (SDHC cards)

**Mission** :

Telescope : Schmidt Cassegrain (300mmφ),

Observation camera : Digital CCD(x2) UV, NIR

# Observation System



## Telescope

Diameter of the Mirror : 300mmφ

Type : Schmidt Cassegrain

Focal Length : 6096mm (with a barlow lens)

Field of View : 158 x 122 arcsec

Control : Direction (Az, El)

4ch Photo  
Multiplier

Tip Tilt Mirror

CCD Image  
Sensor

Image Sensor : Digital 10bit, 659(H)x494(V) pixels

Resolution : 0.2 arcsec (  $5.6 \times 10^{-5}$  deg )

Wave Length : UV(300-450nm), NIR(750-1200 nm)

Data Storage : SDHC memory cards



# Pointing Control System

---

Goal : To control telescope direction toward the target, and to pointing with  
0.1 arcsec accuracy

## The three stage pointing control system

### **First : Gondola Attitude Control**

Attitude of the gondola (Azimuth direction) is controlled to point the solar cell panel toward the sun direction, and keep stable.

### **Second : Coarse Pointing Control**

Azimuth and elevation angle of the telescope are controlled to enter the target in the field of view and keep position of target on the center of the field of view.

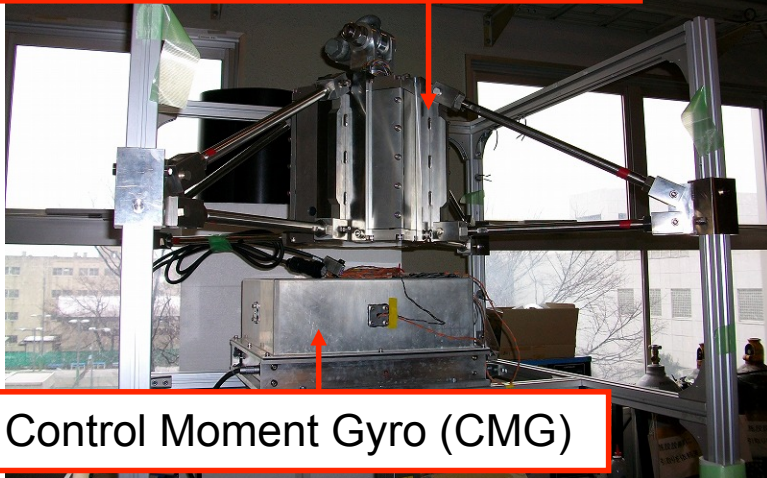
### **Third : Fine Pointing Control**

Position of the target image is controlled with high accuracy during the exposure of the observation camera.

# Pointing Control System

## Gondola Attitude Control

Active Decoupling Motor (DCP)



Control Moment Gyro (CMG)

DCP twists the hanging rope and rotates the gondola.

CMG generates the torque to cancel the vibration.

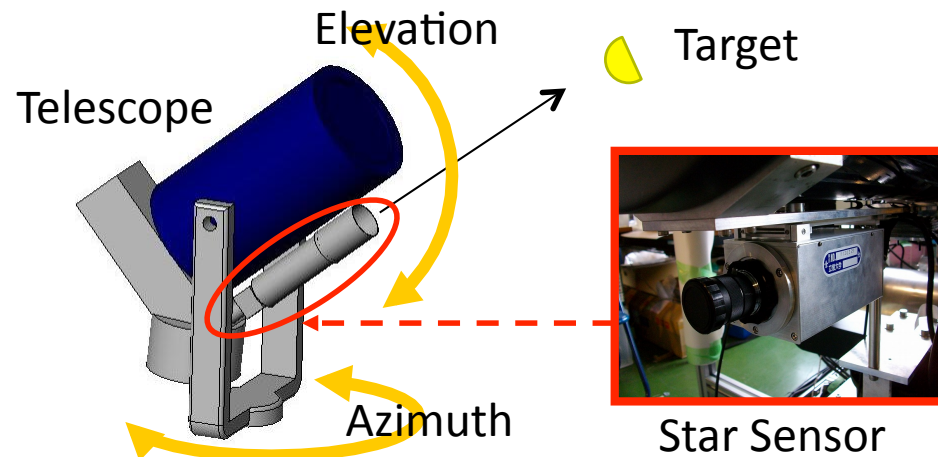
Reference : 2 Sun Sensors (Digital CCD)

- Field of View : 69.8(H) x 55.2(V) deg

- Resolution : 0.11 deg

Required accuracy : < 1.4 deg

## Coarse Pointing Control



Telescope motor controls azimuth and elevation angle.

Reference : Target Star Sensor (Digital CCD)

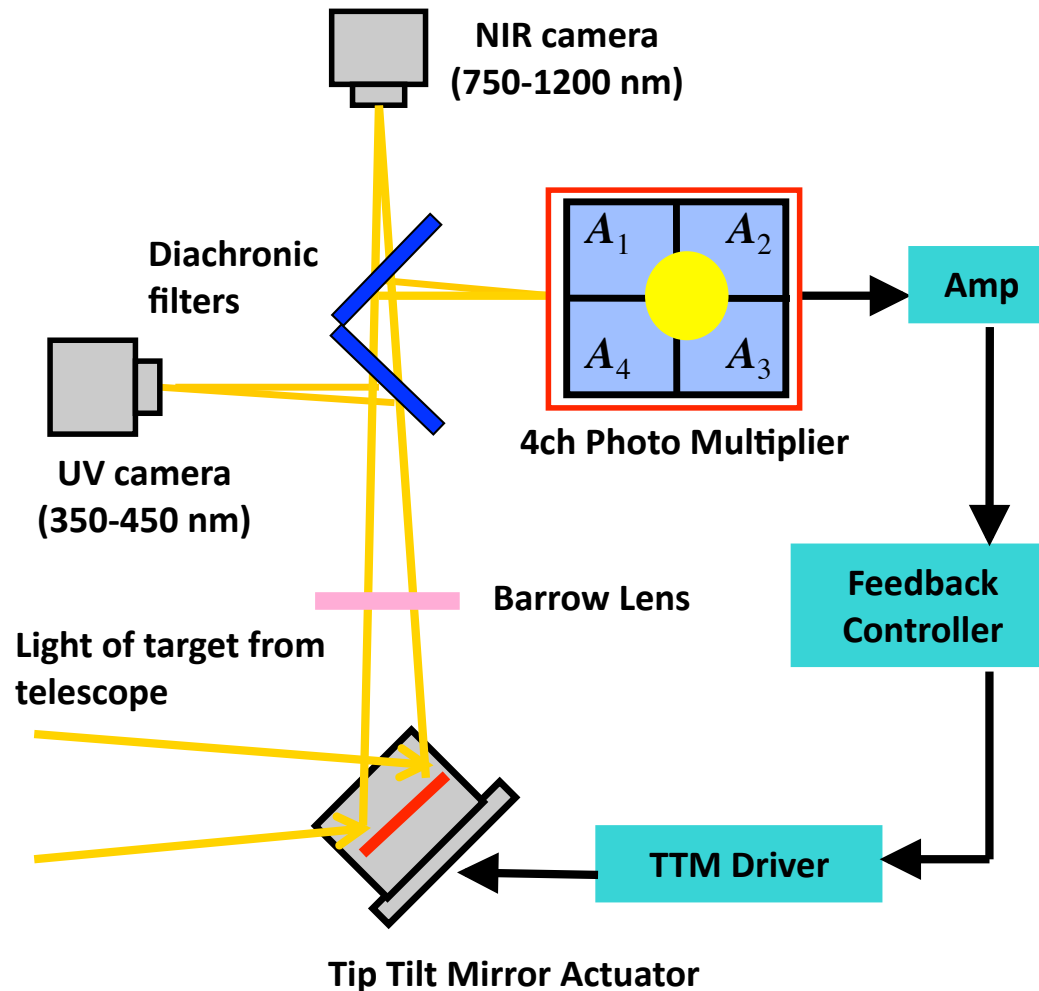
- Field of View : 2.8(H) x 2.1(V) deg

- Resolution : 5.04 arcsec (0.0014 deg)

Required accuracy : < 61.2 arcsec (0.017 deg)



# Pointing Control System



## Fine Pointing Control

Position of target image in FOV of mission cameras is detected and controlled in fine resolution.

Actuator : Tip Tilt Mirror (TTM)  
- Piezo driven bi-axial actuator

Reference : Photo Multiplier Tube (PMT)  
- 4 photo multiplier tube are contained.  
- Each cell outputs electric current by amount of the target light.

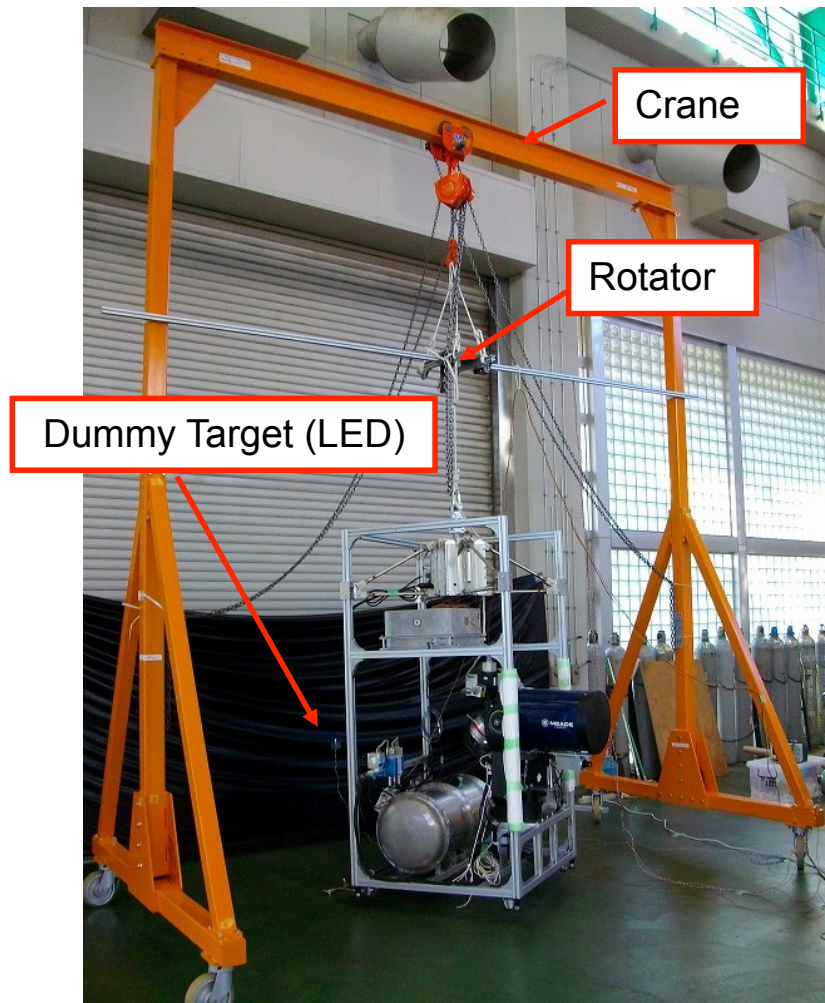
Feedback Controller :  
- Detect the position of the target from PMT outputs, Calculate the control output value to the TTM driver.  
- Control Frequency : 1 kHz

Required accuracy : 0.1arcsec( $2.8 \times 10^{-5}$ deg)

## Overview of Fine Pointing Control system

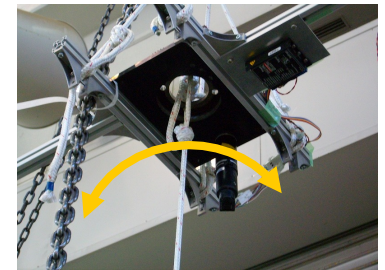
# Ground Test of the Pointing Control System

## Configuration to simulate the flight environment for ground test of unification of gondola attitude control and coarse pointing control



The gondola is hung from the crane with rope.

The rope can be rotated with arbitrary angular velocity or direction by a rotator to simulate the disturbance rotation under flight environment.

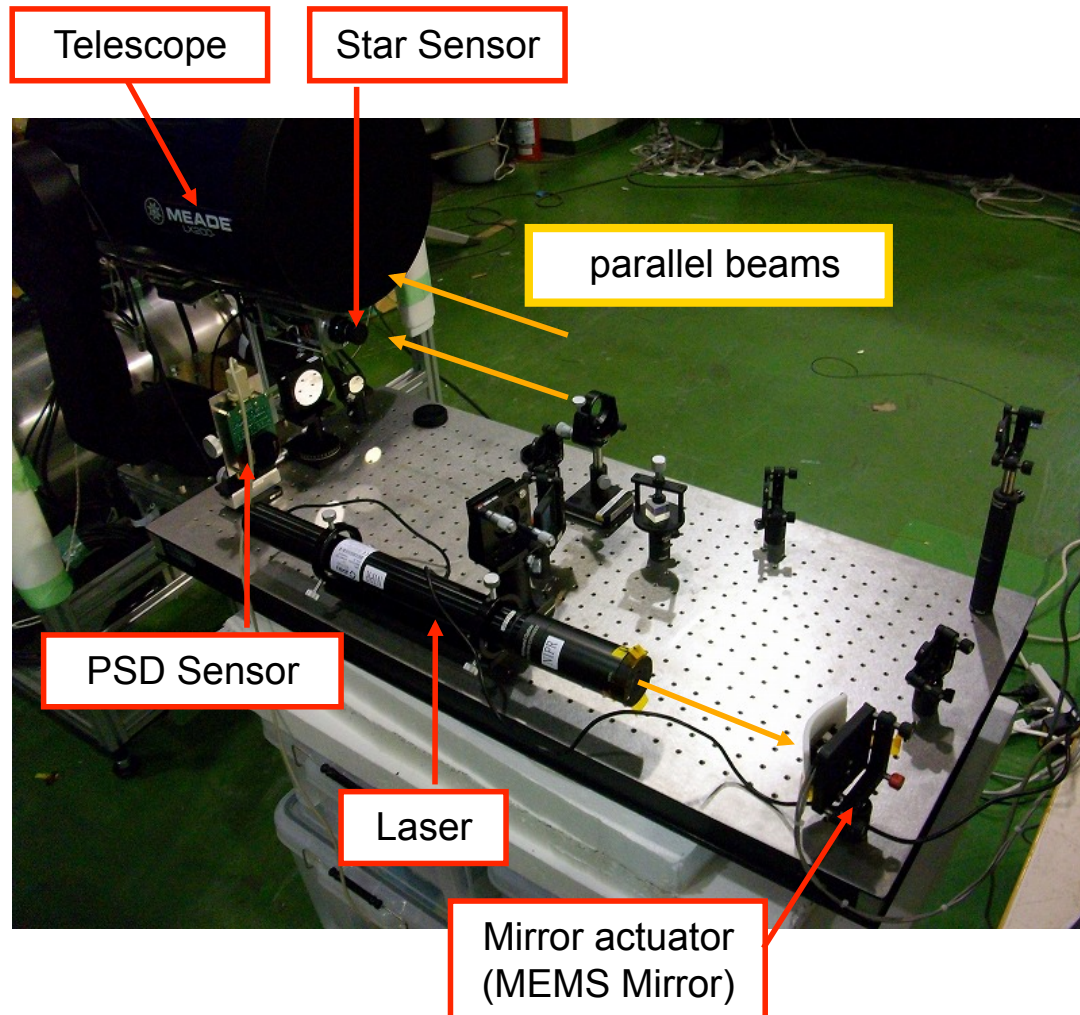


Some LEDs to simulate are attached on the black curtain as dummy light instead of real sun and target.

The performance of gondola attitude control and coarse pointing control is verified, and the control gain is tuned.

# Ground Test of Pointing Control System

## Overview of test bed for performance test of fine pointing control with vibration of target image as disturbances



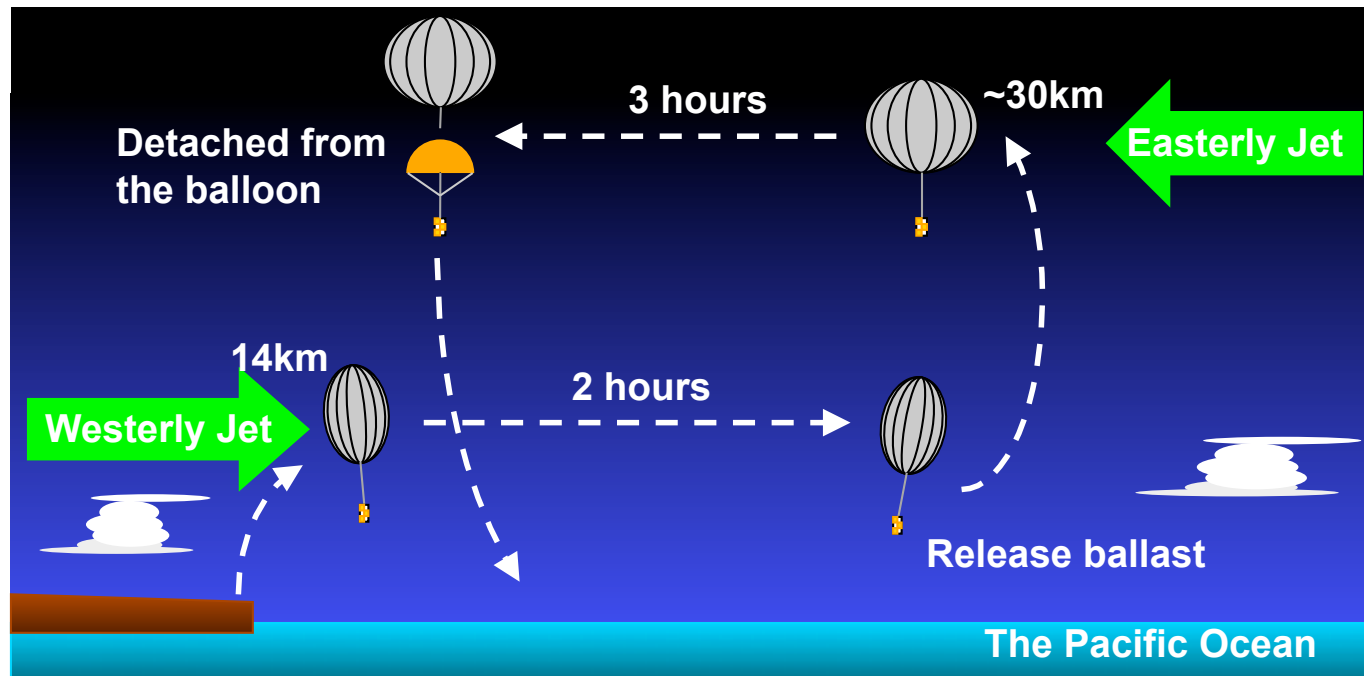
Two parallel beams are generated, and inserted into the telescope and star sensor.

The incident angle of the laser beams is varied with the range of approx. 20 arcsec with a biaxial gimbal mirror build by the MEMS technology.

Current angle of laser is detected by PSD sensor. Position of the light in FOV of telescope is controlled.

The performance of fine pointing control is checked, and the control gain is adjusted.

# Flight Operation in Japan

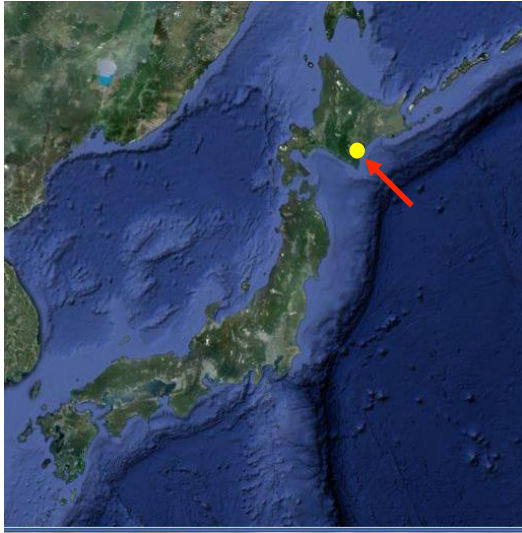


Concept of the Boomerang Flight



# Flight Operation in Japan

---



## **First Flight experiment : on June 3<sup>rd</sup> , 2009.**

- Gondola Subsystems (power supply, house keeping data, ) worked correctly, and some basic data are obtained.

- We couldn't conduct three stage pointing control operation and optical observation because of a trouble occurred on the onboard computer.



## **Second Flight experiment : Summer, 2012.**

- Design of second flight model is based on first flight model, onboard computer is improved.

**Taiki Aerospace Research Field (JAXA)**

# **Possibility of payload contribution?**

Akatsuki payloads

1 mu camera, 2 um camera, 10 mu camera  
UV camera, Lightning/visible photometer

Budget problem

limited budget for Akatsuki team

Schedule overlapping

our launch at ESRANGE will be 2013 summer